Forced tropical motions

Ocean response to wind stress forcing: Yoshida Jet, Stommel model, wave response

Atmospheric response to diabatic heating: Gill model

Parameterization of the atmospheric heating as a function of surface conditions: Gill-Zebiak, Lindzen-Nigam

Thermodynamics

Simple coupled model

The Yoshida Jet:

Consider and interior ocean independent of x and forced by a constant zonal wind stress

$$(\partial_t + \epsilon)u - yv = \tau_x$$

$$(\partial_t + \epsilon)v + yu + \partial_y \eta = 0$$

$$(\partial_t + \epsilon)\eta + \partial_y v = 0$$

The solution is an equatorial jet, in geostrophic balance in meridional direction, and linearly accelerating in time. The meridional velocity is constant, and away from the equator merges to the Ekman transport.

$$v = -\tau_x V$$
$$u = (1 - y V) \tau_x t$$
$$\eta = -\tau_x V t$$

The solution is one valid until zonally dependent arrive from the edges of the wind stress perturbation or from the boundaries of the basin.

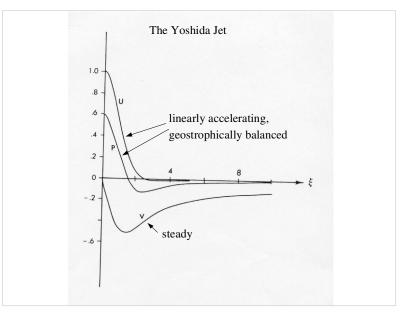
Forced tropical motions

$$(\partial_t + \epsilon)u - yv + \partial_y \eta = \tau_x$$
$$(\partial_t + \epsilon)v + yu + \partial_y \eta = \tau_y$$
$$(\partial_t + \epsilon)\eta + \partial_x u + \partial_y v = -Q$$

Forcing includes wind stress and diabatic heating, dissipation is parameterized as Rayleigh friction and Newtonian cooling. Replace the rate of change with the combination or rate of change and dissipation, and the algebra remains as discussed previously.

(2, 1)

The amplitude of the forcing results from the projection of the stress or diabatic heating onto vertical modes. It is therefore sensitive to the mixing parameterization. As the simplest approach the wind stress forcing applies as a body force over the mixed layer, the diabatic heating represents deep convection and is assumed to project onto the leading mode.

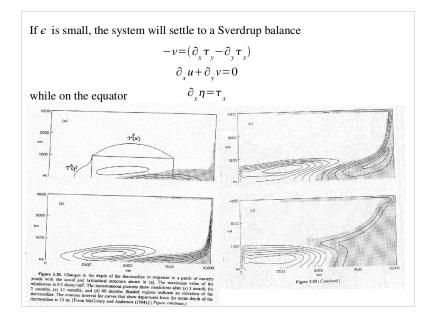


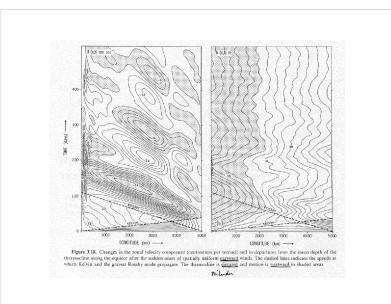
Oceanic response to wind stress forcing

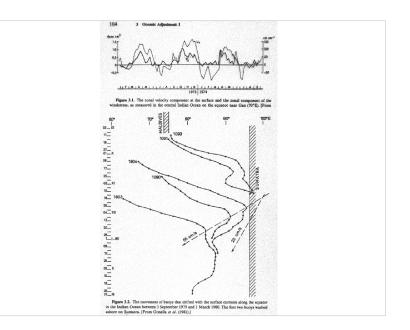
Consider the spin-up problem, switch on the zonal wind stress at initial time, when flow and thermocline depth anomalies are zero (long wave limit)

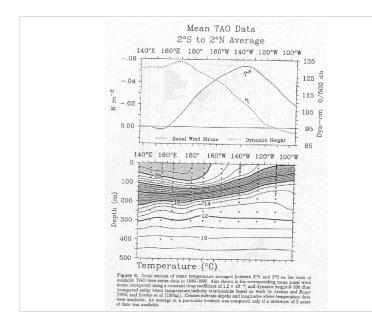
$$(\partial_t + \epsilon)u - yv + \partial_x \eta = \tau_x$$
$$yu + \partial_y \eta = 0$$
$$(\partial_t + \epsilon)\eta + \partial_x u + \partial_y v = 0$$

The solution consists of the particular (steady) response to wind stress and a superposition of equatorial waves such that the initial condition is satisfied. The waves communicate the wind perturbation to the east (Kelvin wave) and west (Rossby waves). Upon reaching the coast the Kevlin wave spread poleward and shed westward propagating Rossby waves, while at the western coast the Rossby wave is converted to a Kelvin wave (experiencing some loss of energy due to the filtered short Rossby waves). These waves then move back into the forcing region and establish the pressure gradients that balance the wind stress.





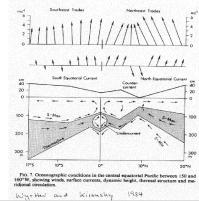




Ekman dynamics - extension to the equator

The steady responses to a (zonal) wind stress in the inviscid limit of the $1\frac{1}{2}$ layer model considered so far include no flow at the equator and no equatorial Undercurrent. This is at odds with the observed currents and temperature structure at the equator that includes a strong shear.

The vertical resolution of the model has to be increased, and the assumption of the wind stress acting as a body force over the entire active layer has to be relaxed. A simple prototype for the effect of mixing is the Stommel model (Stommel, 1960, 'Wind drift near the equator', Deep Sea Res., 6, 298-302).



Consider a layer of constant depth, constant density, with mixing parameterized by a constant eddy exchange coefficient. The layer is forced at the surface by a wind stress, and has a free slip condition at its bottom.

$$\partial_{t}\vec{u} + f\vec{k} \times \vec{u} + \rho_{0}^{-1}\vec{\nabla} p = \partial_{z}A\partial_{z}\vec{u}$$
$$A\partial_{z}\vec{u} = \rho_{0}^{-1}\vec{\tau} , z = 0$$
$$A\partial_{z}\vec{u} = 0 , z = -H$$

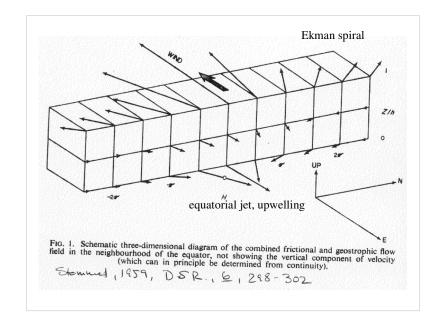
the vertical average of these equations is $\partial_t \vec{u} + f \vec{k} \times \vec{u} + \rho_0 \vec{\nabla} p = (\rho_0 H)^{-1} \vec{\tau}$ while the vertical deviations from the vertical average $\vec{u}' = \vec{u} - \vec{u}$ follow

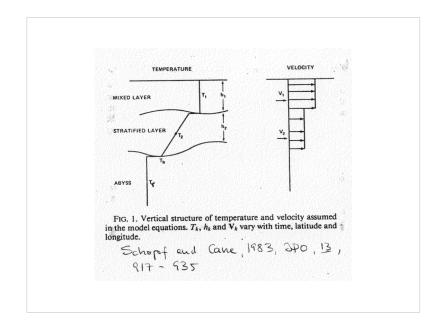
$$\partial_t \vec{u}' + f \vec{k} \times \vec{u}' = \partial_t A \partial_t \vec{u}' - (\rho_0 H)^{-1} \vec{\tau}$$

Note that the pressure term dropped out (since $\rho = const$), and a local problem arises. (The continuity equations decoupled as well: $\vec{\nabla} \cdot \vec{u} = 0$, $\vec{\nabla} \cdot \vec{u} = 0$)

The solution merges the off-equatorial Ekman spiral with a down-wind on the equator and equatorial upwelling.

A two layer version of this model is implemented in the Zebiak and Cane model that will feature prominently in the discussion of ENSO.





Forced atmospheric circulation

Interpret Q as convective latent heating that efficiently projects onto the first baroclinic mode.

Assume long wave approximation and steady state response. Nondimensionalize equations by Rossby Radius L, L/c $c^{1/2}(2\beta)^{-1/2}, (2\beta c)^{-1/2}$ $\epsilon u - \frac{1}{2}yv = -\partial_x p$

 $\epsilon u - \frac{1}{2}yv = -\partial_x p$ $\frac{1}{2}yu = -\partial_y p$ $\epsilon p + \partial_y u + \partial_y v = -Q$

Note that in the steady limit the flow field is divergent. With the transformation q = p + u and the expansion in Parabolic Cylinder functions D_i with the r = p - u recursion relation $d_y D_n + \frac{y}{2} y D_n = n D_{n-1}$

 $d_y D_n - \frac{y}{2} y D_n = -D_{n+1}$

Gill, 1980, Quart. J. R. Met. Soc., 106, 447-462

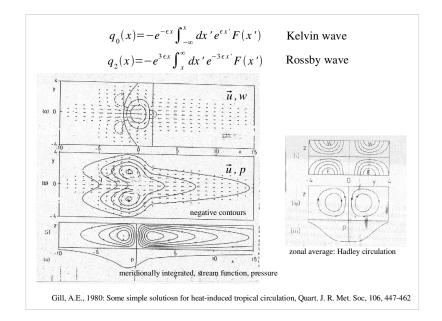
and yields $\begin{aligned} \epsilon q_0 + d_x q_0 = -Q_0 \\ \epsilon q_{n+1} + d_x q_{n+1} - v_n = -Q_{n+1} , n \ge 0 \\ \epsilon r_{n-1} - d_x r_{n-1} + n v_n = -Q_{n+1} , n \ge 1 \\ q_1 = 0 \\ r_{n-1} = (n+1)q_{n+1} , n \ge 1 \\ \end{aligned}$ Example of solution with $\begin{aligned} Q = Q_0 = F(x)D_0 = F(x)e^{-(y/2)^2} \\ \epsilon q_0 + \partial_x q_0 = -F \\ v_0 = 0 \\ \end{cases}$ forced Kelvin wave $\begin{aligned} v_0 = 0 \\ 3\epsilon q_2 - \partial_x q_2 = -F \\ r_0 = 2q_2 \\ v_1 = \epsilon q_2 + \partial_x q_2 \\ q_1 = r_1 = 0 , r_2 = v_2 = 0 \dots \end{aligned}$

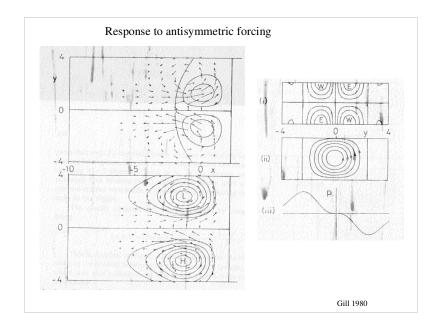
$$q_{0}(x) = -e^{-\epsilon x} \int_{-\infty}^{x} dx' e^{\epsilon x'} F(x')$$
 Kelvin wave
$$q_{2}(x) = -e^{3\epsilon x} \int_{-\infty}^{\infty} dx' e^{-3\epsilon x'} F(x')$$
 Rossby wave

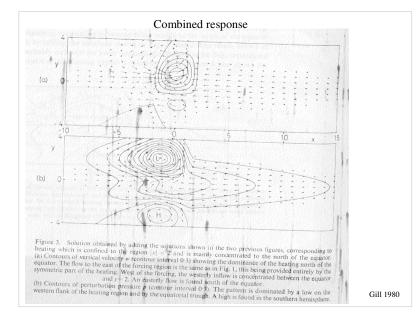
Kelvin wave: no response west of the forcing region, response decays with scale $\varepsilon^{\text{-1}}$ to the east, meridional structure is Gaussian.

Rossby wave: no response east of the forcing region, response decays with $(3\varepsilon)^{-1}$ to the west, meridional structure is D₁.

In response to heating there is poleward flow in the boundary layer due to vortex stretching.





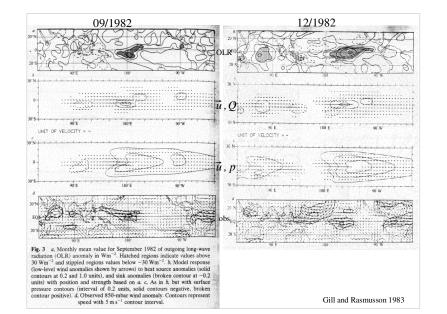


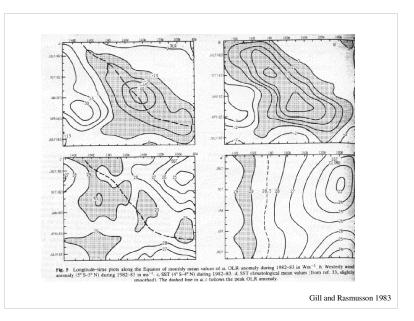
How to specify Q?

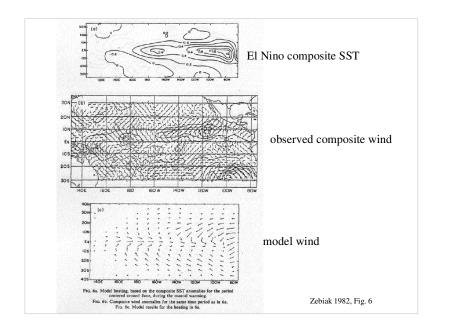
Gill 1980: interpreted Q as latent heat release in deep convection. How can the latent heat release be simulated as a function of surface conditions?

ENSO simulations of Gill and Rasmusson (1983), Nature, 306, 229-234 approximated Q as a function of OLR, and noted a relationship of OLR with migration of the region of highest sea surface temperatures.

Zebiak 1982, JAS, 39, 2017-2027 assumed Q=Q(SST), and based the formalism on local evaporation. Model resemble observations in the western and central equatorial Pacific, but fails in the eastern Pacific and South Pacific.







But latent heat release (precipitation) is not only a function of SST but also of the moisture transport in the boundary layer. Zebiak 1986, Mon. Wea. Rev., 14, 1263-1271

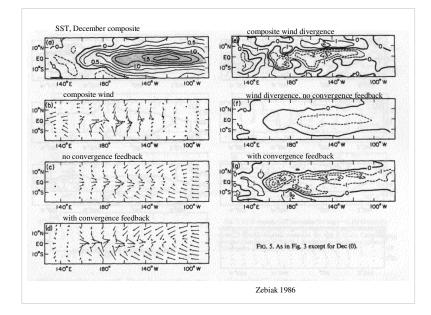
$$Q = Q_{evap}(\overline{T}, T, RH) + Q_{conv}(\overline{\vec{u}}, \vec{u}, \overline{q})$$
$$= \Gamma T' - \beta_{\perp} \nabla \cdot \vec{u}$$

 β_z is nonzero only if total wind field is convergent.

Note that the low level convergence depends on Q and affects Q and leads to an interative solution.

Further refinements such as full moisture budget have been explored (consider the moisture convergence, not just the mass convergence).

Zebiak 1986, Fig. 5



Problems with the Gill-Zebiak model

assumes that the latent heat release in the troposphere affects flow in the boundary layer.

implies very large, local latent heat releases (and gains as a function of SST)

upper level vorticity budget is nonlinear, yet the model is linear the phase speed of the first baroclinic mode is fast, so that a strong damping is needed to constrain the spatial extent of the atmospheric response.

Model type	Model source	Mechanical damping $\epsilon_{\rm D}^{-1}$ (days)	Venting or thermal damping ϵ_{T}^{-1}	Wave speed $m s^{-1}$	Typical forcing (W m ⁻²)
Convective (full troposphere)	Gill-Zebiak	1-2	1-2 days	60	75
	observations	≈10	≈15 days	≈40	15-20 (evaporative)
Boundary layer	Lindzen–Nigam Reduced gravity	≈1-2	30 min	≈16*	5-10 (virtual sensible)
	(section 2b)	$\approx 1-2$	8 h	≈ 18	5-10 (virtual sensible)
	Observations	≈1-2	≈3–12 h	≈18	5-10 (virtual sensible

Lindzen and Nigam, 1987, JAS, 44, 2418-2436

considers boundary layer only

SST anomalies produce sensible and latent heat fluxes and affect the virtual temperature, and therefore pressure in the boundary layer the temperature anomaly in the boundary layer is

$$T_{a}'(x, y, z) = T_{y}'(x, y)(1 - z\gamma/H_{b})$$

where H_b is the boundary layer height (700 mb height surface).

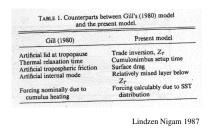
The momentum equation in the boundary layer is (the true and tried)

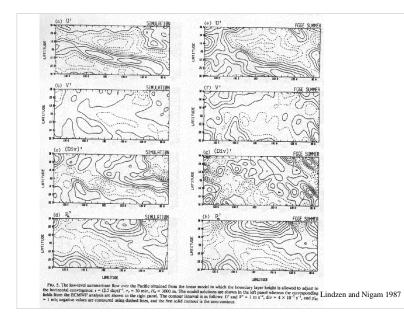
$$\epsilon_{D}\vec{u} + f\vec{k} \times \vec{u} + \vec{\nabla}\phi = 0$$

Pressure is affected by temperature changes and by the fluctuating top of the boundary layer $\phi = g(1 + \alpha H_b/\overline{T}_c)h + \Gamma T_u'$

which depends on the balance of the mass convergence in the boundary layer and entrainment, acting on a convective adjustment time $-H \vec{\nabla} \cdot \vec{u} = \epsilon_{h} h$
$$\begin{split} \epsilon_{_D} \vec{u} + f \vec{k} \times \vec{u} + \vec{\nabla} \phi = 0 \\ \epsilon_{_D} \phi + c_{_{\ln}}^2 \vec{\nabla} \cdot \vec{u} = -\epsilon_{_D} \Gamma_{_{\ln}} T_{_V} ' \\ c_{_{\ln}}^2 = gh \frac{\epsilon_{_D}}{\epsilon_{_T}} \end{split}$$

Thus, the equations of the Lindzen and Nigam model are very similar to the Gill model, but the physical interpretation and parameters are very different.





Reduced gravity model: Battisti, Sarachik, Hirst, , 12, 2956-2964

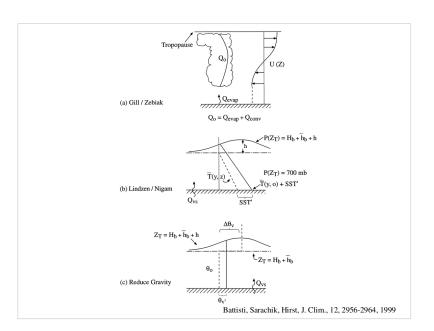
This model is similar to Lkindzen and Nigam, but the virtual temperature perturbations are well mixed in the boundary layer. The boundary layer is capped by a constant inversion. The free atmosphere acts to reduce gravity and to absorb mass and heat without significant flow above the boundary layer (reduced gravity). Damping time of the boundary layer depends on its mass convergence: If convergence is diagnosed, the boundary layer height is damped rapidly (8 hrs, convection), else a slower entrainment rate of 1-2 days is assumed.

$$\epsilon_{D}\vec{u} + f\vec{k} \times \vec{u} + \vec{\nabla}\phi = 0$$

$$\epsilon_{D}\phi + g'H(1-\beta')\vec{\nabla} \cdot \vec{u} = -\epsilon_{D}\Gamma T'$$

$$\beta' = \frac{1 - \frac{\epsilon_{d}}{\epsilon_{m}}}{6} \text{ if convergent}$$

$$0 \text{ else}$$
Form of model is again similar to Gill's model.

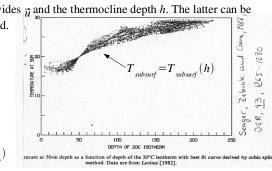


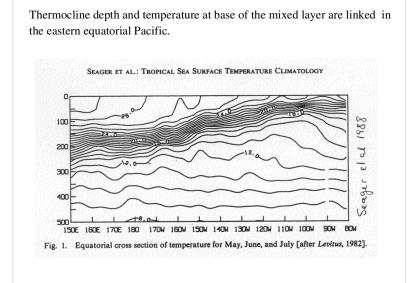
To couple the ocean and atmosphere models, we need to consider SST. In the spirit of the reduced (simple) models discussed so far, consider the temperature balance of a surface mixed-layer of depth H_1 and vertically constant temperature T $\partial_{t}T + \vec{u} \cdot \vec{\nabla}T = H_{1}^{-1} \Big[(\rho_{0} c_{p})^{-1} Q_{AS} - w (T - T_{subsurf}) \Big]$

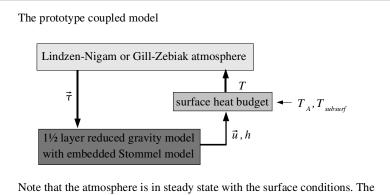
The ocean model provides $\vec{\mu}$ and the thermocline depth h. The latter can be empirically determined.

The air-sea heat flux Q_{AS} is expressed in terms of surface temperature T and equilibrium temperature T_A $Q_{AS}(T) = -\alpha(T - T_A)$

Thermodynamics of equatorial SST







Note that the atmosphere is in steady state with the surface conditions. The ocean provides the memory and slow evolution due to thermal inertia of the mixed layer, and more importantly due to Rossby and Kelvin wave dynamics.

This model forms the basis for discussions of the coupled dynamics in the tropics