



2.5m

J. D. Woods (1968)



Grae Worcester, University of Cambridge



Vanja Zecevic, University of Sydney



Lewis Fry Richardson

Richardson Number

Necessary condition for instability:

$$R_i\left(=\frac{N^2}{S^2}\right) \le 0.25$$

somewhere in the flow Miles (1961), Howard (1961)

$$N = \sqrt{-\frac{g}{\rho_o} \frac{\partial \rho(z)}{\partial z}}$$

Thorpe (1987, JGR 92, 15, 5231-5248

Thorpe: Transitional Phenomena of Turbulence in Stratified Fluids



Fig. 7. Stages in the transition to turbulence of Kelvin-Helmholtz instability.





Figure 1. Differenced brightness images from H20 at -90-s intervals showing emerging vorticity dynamics as KH billows that are deformed and/or misaligned evolve to larger amplitudes and exhibit various instability dynamics in an -50×50 -km field of view (with north at top and east at left). See the image dimensions at upper left and notation and highlights of the relevant features where they first arise and evolve thereafter.



Fritts et al 2020

Figure 6. (a–f) KHI tube and knot evolution due to linking of misaligned billows spanning $\sim 2 T_b$ viewed from above beginning at the initial time in Figures 3 and 4. Shown are volumetric views of A_2 , with small values dark green and large values red in a subset of the full domain. Here, x and y increase toward top and left, respectively, and positive mean ζ_p is toward increasing y (to the left). Panel (a) shows the axes.

Observations: brightness

Model: rotation



Jørgen Holmboe

Holmboe instability



Smyth and Winters JPO 2003



Mixing in models of the Ocean and Atmosphere



Turbulent fluxes

$$-\overline{u'w'} = \kappa_m \frac{\partial \overline{u}}{\partial z}$$

$$-\overline{w'T'} = \kappa_h \frac{\partial \overline{T}}{\partial z}$$

Scaling:

$$\epsilon = \ell_v^2 N^3 f(Ri)$$

Note: if f(Ri)=1 then $\ell_v = \ell_o = v(\epsilon/N^3)$, the Osmidov scale, and if $f(Ri)=Ri^{-3/2}$ (as $Ri \rightarrow 0$) then $\ell_v = \ell_c = v(\epsilon/S^3)$, the Corrsin scale.

$$\kappa = \gamma \ell_v^2 N f(Ri)$$

See Ijichi and Hibiya JPO(2018) for estimates of γ

$$\kappa = \gamma rac{\epsilon}{N^2}$$

Osborn 1982







Salinity along 165E: April 1988











https://www.pmel.noaa.gov/tao/drupal/disdel/







"Big whirls have little whirls that feed on their velocity, and little whirls have lesser whirls and so on to viscosity" Lewis Fry Richardson





Eq, 156E







Filtered









The variation of $\varepsilon \sim N$ for constant Ri has implications for the scaling of the turbulence

$$\epsilon = \ell_v^2 N^3 f(Ri)$$

then

 $\kappa = \frac{\gamma \epsilon}{N^2}$

$$\ell_v = c \frac{u_t}{N}$$
$$u_t \simeq 0.1 \tilde{u}$$





Observed versus Estimated ε



1N, 156E

Deconstructed $\epsilon \sim \tilde{u}^2 N f(Ri)$



Eq, 156E



1N, 156E



Eq, 156E

- Important to resolve flow features generating turbulence
- Need to consider turbulent length scale
- What do you do if relevant scales are not resolved?

Parameterization if S², N² NOT resolved

$$\kappa(\mathbf{x},t) = \frac{\gamma}{N^2} \ \epsilon(S^2,N^2)$$

 $(S^2, N^2) \sim (\langle U \rangle, \langle N \rangle^2, F(x - x', t - t'), F_T \downarrow)$

Instabilities

Shear + Oscillatory flow = Inertial Instability and/or PSI

producing small vertical scales



Model developed by Hidenori Aiki



FIG. 2. Snapshots of the meridional component of velocity on days 40 (top), 240 (middle), and 480 (bottom row). Left panels: 2d; right panels: 3d at a given longitude. Vertical dashed lines mark the boundaries of the region of initial anomalous background vorticity.

Persistent presence of small vertical scale velocity features during three-dimensional equilibration of equatorial inertial instability

Natarov and Richards (2015)



FIG. 10. Evolution of vertically and zonally averaged vertical eddy diffusion coefficient: (a) 2d, (b) 3d.

Wind

Production of near-inertial oscillations

Wind impulse applied to mixed layer at 5N (Linear continuously stratified model)





Linear model forced with QuikSCAT along 156E





Figure 4. Base-10 logarithm of time-averaged vertical eddy viscosity coefficient below the mixed layer. (upper panel) Period I (equatorial winds). (lower panel) Period II (off-equatorial winds).

On-Eq winds

Off-Eq winds

Enhanced Energy Dissipation in the Equatorial Pycnocline by Wind-Induced Internal Wave Activity

Natarov and Richards 2019

Too hard? Well, try increasing *vertical* resolution



FK: R/V Falkor cruise, August 2015 at 170°W, 1°N

model high: 1/3° horizonal, 3 m vertical (top 402 m)

model low: 1/3° horizonal, 5-37 m vertical (top ~400 m) (16 m average)

1/27 (model): 1/27° horizonal, 3 m vertical (top 402 m) ΔSST



Model high – Model low